

## SUBMILLIMETER-WAVE ROTATIONAL SPECTROSCOPY OF $\text{H}_2\text{F}^+$

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Five pure rotational transitions of  $\text{H}_2\text{F}^+$  were observed in the 473-774 GHz range with a backward-wave oscillator based submillimeter-wave spectrometer<sup>a</sup>. The  $\text{H}_2\text{F}^+$  ion was generated in an extended negative glow discharge in a gas mixture of hydrogen fluoride generated by heating potassium hydrogen fluoride ( $\text{HF}_2\text{K}$ ) granular powder at 150°C-160°C and hydrogen in an argon buffer. A simultaneous analysis of the rotational lines with 120 combination differences for the ground state derived from the infrared spectra obtained by Schäfer and Saykally<sup>b</sup> and Fujimori et al.<sup>c</sup> was carried out to determine the precise molecular constants for the ground state. The rotational transition frequencies that lie below 2 THz were calculated, together with their estimated uncertainties, to facilitate future astronomical identifications. Recently  $\text{H}_2\text{Cl}^+$  was detected in NGC 6334I and Sgr B2 with the Heterodyne Instrument for Far-Infrared (HIFI) on board the Herschel Space Observatory<sup>d</sup>, and HF was also detected in a wide variety of interstellar clouds with the same facility<sup>e</sup>. The proton affinity of HF is smaller than those of  $\text{N}_2$  and CO, so the abundance of  $\text{H}_2\text{F}^+$  is likely to be low in dense molecular clouds. We will discuss abundances of  $\text{H}_2\text{F}^+$  in diffuse molecular clouds, considering various chemical reaction rates.

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<sup>a</sup>T. Amano and T. Hirao, *J. Mol. Spectrosc.*, **233**, 7 (2005)

<sup>b</sup>E. Schäfer and R. J. Saykally, *J. Chem. Phys.*, **81**, 4189 (1984)

<sup>c</sup>R. Fujimori, Y. Hirata, K. Kawaguchi, and I. Morino, RE02, *The 65<sup>th</sup> OSU International Symposium on Molecular Spectroscopy* (2010)(Columbus, OH)

<sup>d</sup>D. C. Lis et al., *Astron. Astrophys.*, **521**, L9 (2010)

<sup>e</sup>D. A. Neufeld et al., *Astron. Astrophys.*, **518**, L108 (2010); T. G. Phillips et al., *Astron. Astrophys.*, **518**, L109 (2010); P. Sonnentrucker et al., *Astron. Astrophys.*, **521**, L12 (2010)