

NUCLEAR SPIN OF H_3^+ IN DIFFUSE MOLECULAR CLOUDS

KYLE N. CRABTREE, NICK INDRIOLO, HOLGER KRECKEL, BRIAN A. TOM,^a BENJAMIN J. McCALL, *Department of Chemistry, University of Illinois, Urbana, IL 61801, USA.*

In diffuse molecular clouds (environments with high molecular fraction, but low CO abundance), the relative populations of the $J = 0$ (*para*) and $J = 1$ (*ortho*) rotational levels of H_2 are often used as a measure of the cloud kinetic temperature, T_{01} . Typically, T_{01} is on the order of 50-70 K, but in similar environments, the excitation temperature $T(\text{H}_3^+)$ derived from the $(J, K) = (1,0)$ (*ortho*) and $(1,1)$ (*para*) rotational levels of H_3^+ is 20-40 K. We have extended the number of sight lines in which both T_{01} and $T(\text{H}_3^+)$ have been measured from two to five, and in four of the five cases, the two temperatures are discrepant in the same cloud. Using a steady state chemical model based on rate coefficients calculated with a microcanonical statistical approach, we find that the discrepancy between T_{01} and $T(\text{H}_3^+)$ likely arises from incomplete thermalization caused by competition between the thermalization reaction $\text{H}_3^+ + \text{H}_2 \rightarrow \text{H}_2 + \text{H}_3^+$ and dissociative recombination of H_3^+ with electrons.

^aPresent Address: Department of Chemistry, United States Air Force Academy, Air Force Academy, CO 80840, USA