

THE NEAR IR BULK and JET-COOLED  $\text{NO}_2$  SPECTRA by FTS, ICLAS, CRDS and LIF

RÉMY JOST, PATRICK DUPRÉ, ANTOINE DELON, SYLVAIN HEILLIETTE, PATRICE THEULÉ, *Grenoble High Magnetic Field Laboratory, CNRS-MPI, BP 166, 25 Rue des Martyrs 38042 GRENOBLE, Cedex 9, France*; ALAIN CAMPARGUE, GABRIELE WEIRAUCH, *Laboratoire de Spectrométrie Physique, CNRS-Université J. Fourier de Grenoble, BP 87, 38402 SAINT MARTIN D'HÈRES Cedex, France*; JOHANNES ORPHAL, *Laboratoire de Photophysique Moléculaire, CNRS-Université Paris-Sud, Bât. 350, Centre d'Orsay, 91405 ORSAY Cedex, France*; JOHN P. BURROWS, SABINE DREHER, *Suzanne Voigt Institute of Environmental Physics and Remote-Sensing University of Bremen, PO Box 330440, 28334 BREMEN, Germany*.

In order to characterize experimentally the  $\tilde{X}^2A_1 - \tilde{A}^2B_2$  conical intersection in  $\text{NO}_2$  we have extensively studied its near IR spectrum, both in a bulk cell and in a slit jet expansion. The advantages and disadvantages of the four techniques (FTS, ICLAS, CRDS, LIF) will be briefly summarized and their performances compared. Actually, only the slit jet spectra are analyzable, the room temperature spectrum being highly congested. More than 100 vibronic levels belonging to the  ${}^2B_2$  symmetry have been observed by absorption from 10800 to  $13900\text{ cm}^{-1}$  whereas only 15 vibrational levels of the  $\tilde{A}^2B_2$  state (bright levels) are expected in that range. The  $\tilde{X}^2A_1 - \tilde{A}^2B_2$  vibronic interactions strongly mix the (dark) set of  $\tilde{X}^2A_1$  levels (about 85 levels belonging to the  ${}^2B_2$  symmetry) with the set of 15  $\tilde{A}^2B_2$  bright levels. In addition, numerous hot bands have also been observed in that range. Most of these vibronic levels have an irregular rotational structure, even if most of the  $K$ -subbands seem almost regular. This results from the combination of large energy differences between interacting levels and large matrix elements leading to smoothly varying mixing coefficients in each  $K$  stack. The iterative comparison of the band origins, intensities and rotational constants with those predicted by a model Hamiltonian is the clue of the understanding of the  $\text{NO}_2$  conical intersection.